

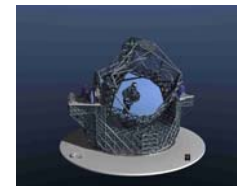
TEOPS: "Groups, Commonality, Activity, Capabilities & Silicate bonding"

- What is TEOPS?
- Technology for Experimental and Observational Physics in Scotland
- Initiative in SUPA Astrophysics and Space Research theme
- "Spans the areas of particle physics, astrophysics and astronomy with a common theme of leading edge technology"
- Collaboration between UK ATC and Glasgow University Institute for Gravitational Research and Experimental Particle Physics groups

Institute for Gravitational Research (IGR)

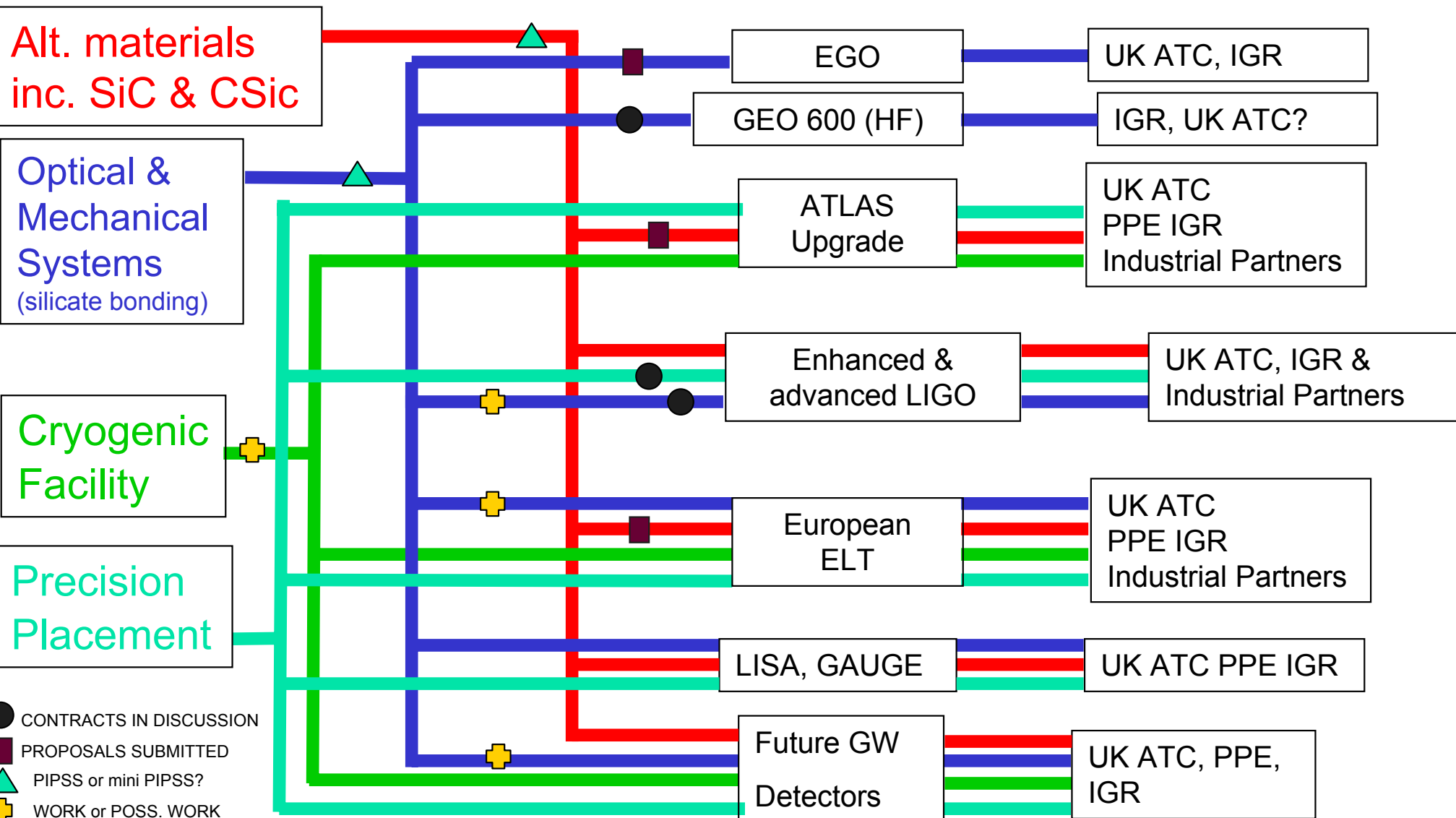
- *The work of the Institute is focused on*
 - *the development of detectors to search for gravitational waves from astrophysical sources*
 - *data analysis activities within the LIGO Scientific Collaboration.*
- *The main areas of experimental research are*
 - *development of precision novel interferometric techniques*
 - *development of systems of ultra low mechanical loss for the suspensions of mirror test masses along with research towards the space-based LISA mission*
 - Development of multiple pendulum systems using silica fibres to support the test masses
 - New bonding technology (hydroxide-catalysis bonding), which exhibits very low mechanical loss and is compatible with ultra-high vacuum
- *The technology developments within the IGR are of broader relevance to a number of areas of current PPARC interest and extensions of the bonding technology are being pursued with general application to precision optical systems on the ground and in space.*

- Experimental Particle Physics - Detector Development group
 - involved in a wide range of projects related to imaging, radiation detection and detector development, within particle physics, medicine, biology
- *Examples of current/recent projects include:*
 - CERN ATLAS - Production and testing of modules for the LHC/ATLAS
 - CERN Medipix - High sensitivity X-ray imaging for medical and synchrotron applications
 - Retinal imaging - Measuring the electrical activity of retinal tissues



- UK Astronomy Technology Centre, Edinburgh
 - **MIRI** - hosting the European PI and opto-mechanical design leads for this key instrument on the JWST, successor to the Hubble Space Telescope
 - Involved in building instrumentation for the most exciting international ground-based and space-borne astronomy projects
 - **WFCAM** - the largest infrared camera ever built, a cryogenic instrument now undertaking unique surveys in the Northern Hemisphere skies
 - **SCUBA2** - the successor to SCUBA, one of the most successful ground-based instruments ever built, utilising a new generation of sub-millimetre CCD-like detectors
 - **European Extremely large telescope (E-ELT)** - The UK ATC is leading the UK's work towards an optical and infrared telescope of up to 42 m in diameter, recently approved into the design phase by ESO Council
- ATC is also involved in several UK and European network and technology development initiatives
- Technology and research strengths complementary to and highly relevant to IGR and PPE groups in Glasgow

AREAS OF COMMONALITY/ INTEREST



TEOPS: Capabilities #1

- VMC (HAAS)
 - 5 axis
- LISA Pathfinder lab
 - precision assembly



- SRDG - Optical
Characterisation Suite
- Detector Characterisation
Suite (PPE)

Optical Characterisation Suite

- Zygo interferometer
 - Suitable for surface flatness and curvature measurements on mirrors
 - Flatness measurements of <1 nm are possible over a 4 or 6 inch wafer
- Wyko NT1100 optical surface profiler
 - Field of view from 3.5×4 mm to 50×50 μm
 - Step heights up to several mm
 - Surface roughness measurements down to 0.5 nm
- TM1000 tabletop SEM
 - ~ 30 nm resolution on insulating substrates as well as conductive (reduced charging)
 - Totally self contained no additional pumps
 - ~ 2 min pump down



Detector Development Lab (PPE)

- Automatic wire-bonding system
 - used for high precision, high density connections between detectors and readout electronics
 - used for the wire-bonding of SCT modules for the ATLAS project at CERN (~1.5 Mbonds)

- Cascade Microtech S300 probe station
 - sets the measurement standard for 300mm on-wafer test
 - applications
 - device characterization and modeling
 - wafer-level reliability
 - design de-bug
 - IC failure analysis
 - has the precision and versatility needed for the most advanced semiconductor processes and aggressively scaled devices
 - allows the group to create the perfect on-wafer measurement environment



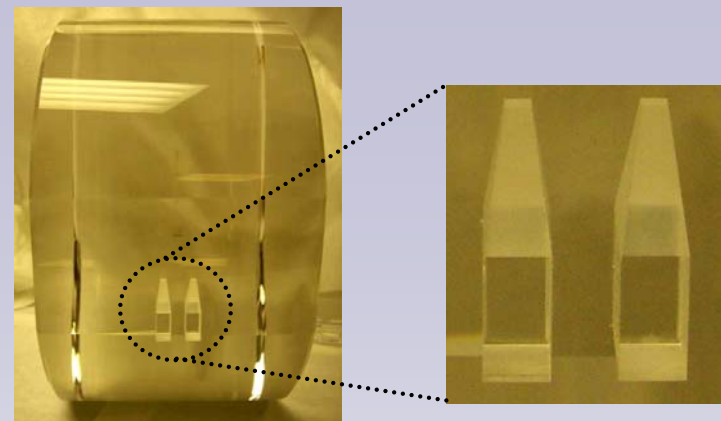
TEOPS: Capabilities #4

- Cryogenic Material Property Test Bed
 - Based at UK ATC

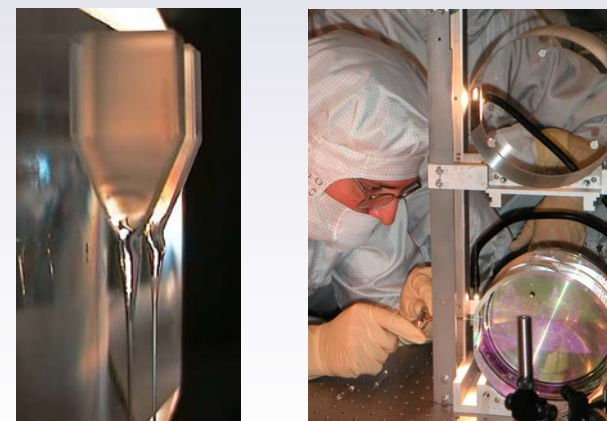


Current applications

- Originally developed for NASA's Gravity Probe B mission, launched April 2004. (Gwo et al.)
- GEO600 currently operates with quasi-monolithic fused silica suspensions and mirrors. This **technology allows improved thermal noise** in the suspension systems.
- Construction of the ultra-rigid, ultra-stable optical benches for the LISA Pathfinder mission.



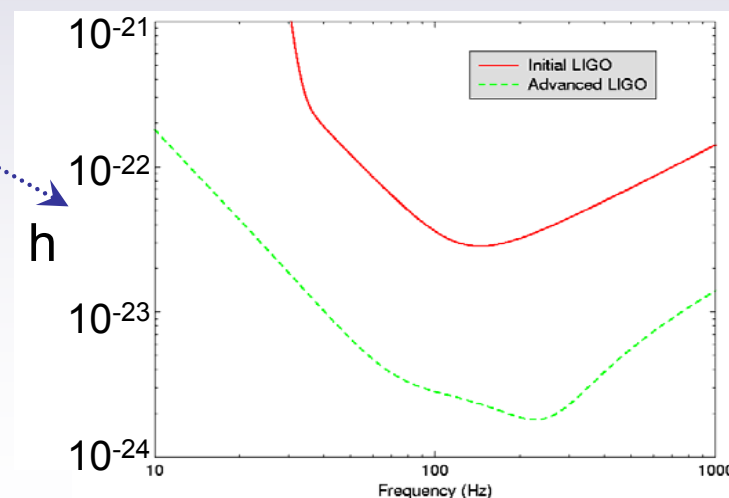
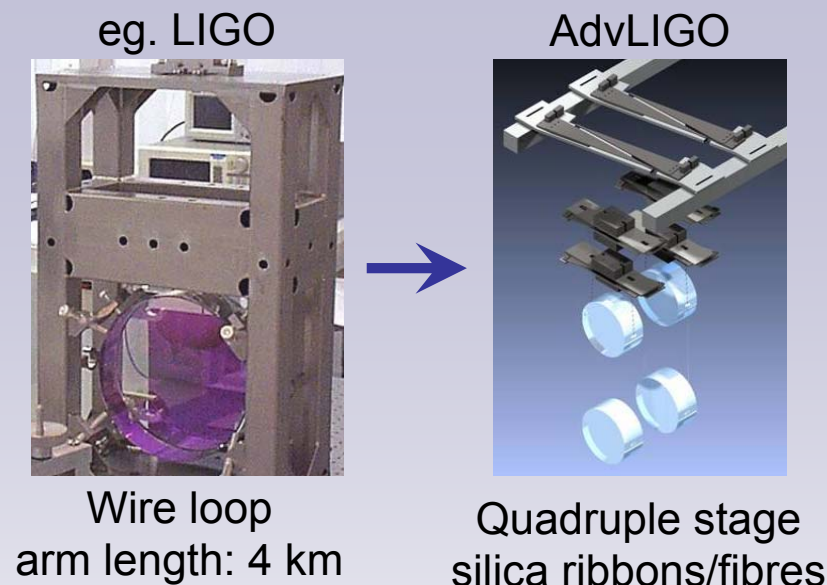
Picture of a GEO600 sized silica test mass in Glasgow with silica ears jointed using hydroxy-catalysis bonding



Silica fibres are welded to the ears in the completion of the lower-stage of the GEO600 mirror suspension.

Planned applications

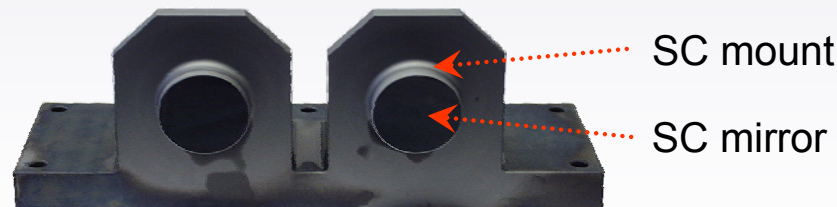
- The planned upgrades for AdvLIGO and Advanced VIRGO plan to incorporate the GEO600 technology for significantly improved thermal noise performance (in addition to other improvements, e.g higher power lasers).
- Construction of the ultra-rigid, ultra-stable optical benches for LISA.



Design sensitivity curves for the LIGO and AdvLIGO detectors.

Motivation

- Many optical systems layouts have stringent requirements for strength, rigidity, stability and **alignment**.
- Hydroxy-catalysis bonding fulfills all these requirement.
- One possible disadvantage of this technique is that the time taken for a typical bond to “set” at room temperature is in the region of a few tens of seconds. This only allows a **short period of time in which to align** the various components on the optical bench
 - Glasgow has investigated how to extend the settling time of hydroxy-catalysis bonds through varying the hydroxide concentration and lowering the temperature.
(Reid et al, Physics Letter A, 2007)



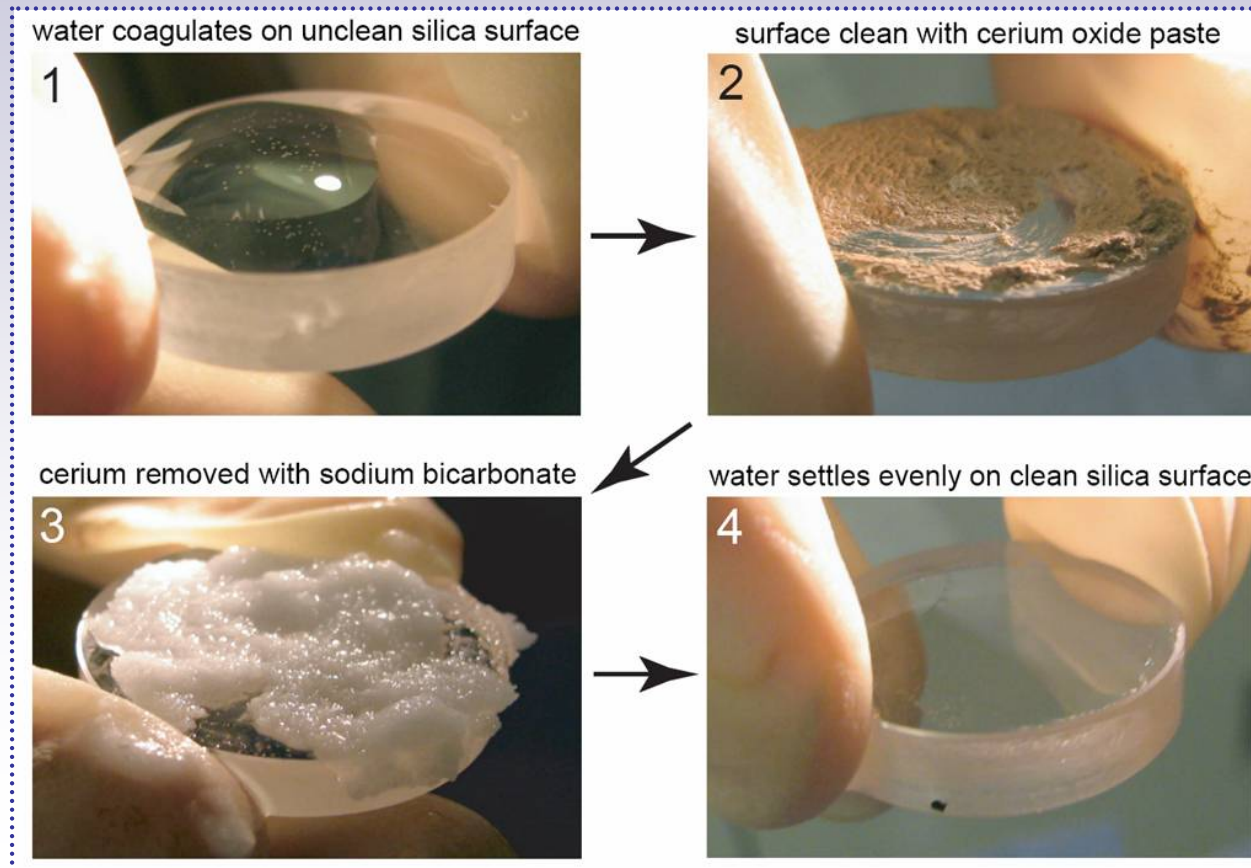
Picture of silicon carbide mirrors bonded to a silicon carbide base mount for GAIA



The processes involved in hydroxy-catalysis bonding

Surface preparation

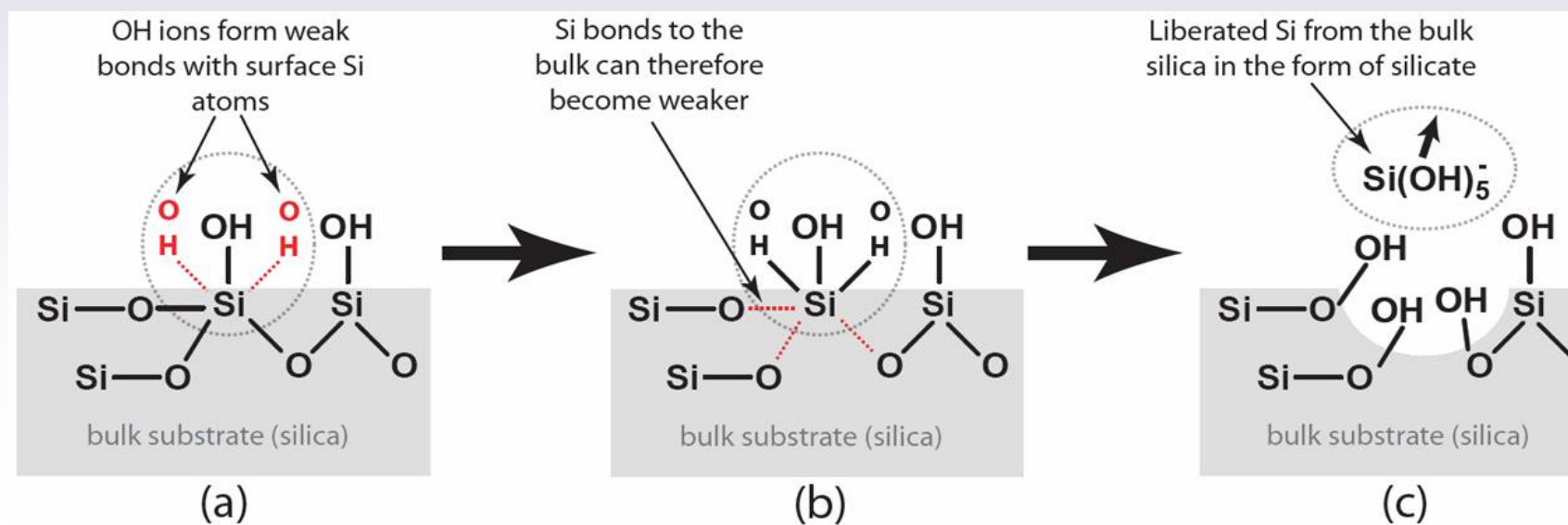
- Contaminants on the silica surface will likely inhibit hydration. The silica surfaces to be jointed are thus taken through a cleaning process to remove any contaminants and to ensure maximum hydration:



The processes involved in hydroxy-catalysis bonding

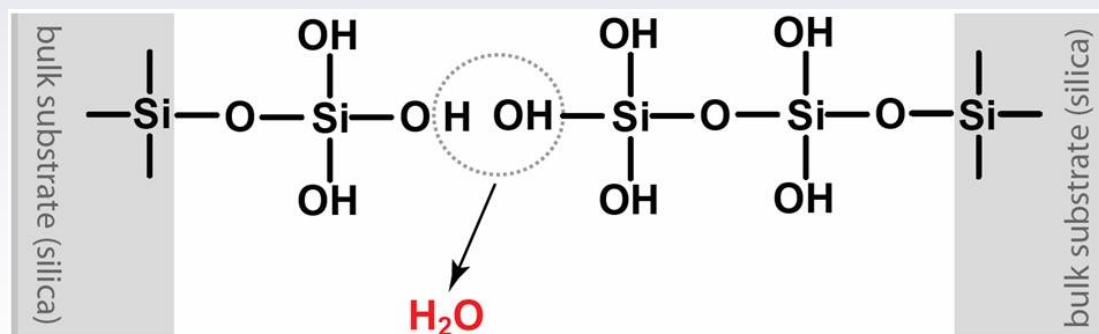
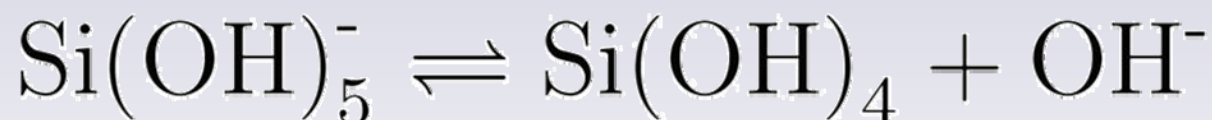
Etching — Etching of the silica surfaces to be bonded

- Placing a solution with a high concentration of OH^- ions on the surface of silica causes etching to take place.
- Free OH^- ions form weak bonds with silicon atoms on the substrate surface causing the original lattice **bonds to weaken**
- It becomes possible for the silicate molecule to break away from the bulk structure, producing **$\text{Si}(\text{OH})_5^-$ molecules in solution**.



Bonding – Polymerisation of silicate in solution

- However, below pH 11, the silicate ion hydrolyses to soluble $\text{Si}(\text{OH})_4$ and OH^- . When the concentration of $\text{Si}(\text{OH})_4$ molecules reaches 1→2%, the solution polymerises and becomes “rigid” (R.K. Iler, 1979, *The Chemistry of Silica*).
- $\text{Si}(\text{OH})_4$ is a **monomer** which likes to form a **polymer** arrangement:



- Bonding starts to dry and evaporates H_2O
- Bond thickness ~ 100 nm



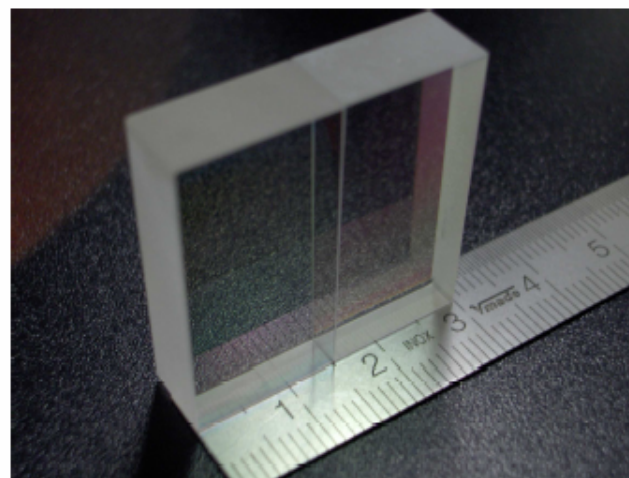
Investigations into hydroxy-catalysis bonding

Development
Theory of silicate bonding
Cleaning process for bonding

Material for silicate bonding
Clean room

De-Bonding

- De-bonding is possible within a few hours from bonding, depending on size, shape and quality of the bond
- Using ultra-sonic bath with detergent solution (10% Decon[®])
- Each bonding and de-bonding process can slightly damage the surface for bonding



Ongoing work in Glasgow includes:

- Further investigation of bonding samples with a **ground finish**.
 - align optics **without danger of optical contacting**
 - **no time constraint** on achieving alignment
 - apply bonding solution with pieces in situ

- Bonding other materials

e.g. Silicon-silicon bonded samples
for verifying the feasibility of monolithic
silicon suspensions for 3rd generation
gravitational wave detectors:

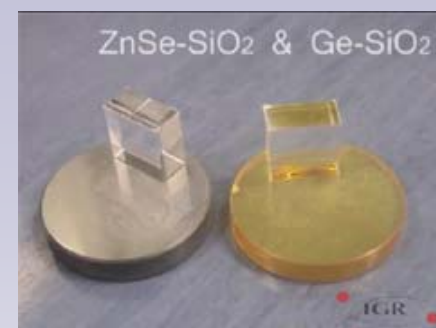
- **mechanical strength**
- **thermal conductivity (at low T)**
- **mechanical loss**



Silicon-silica hydroxy-catalysis bond

Investigations into hydroxy-catalysis bonding

- Successful test for Astrium D for feasibility of ZnSe/SiO_2 and Ge/SiO_2 bonds

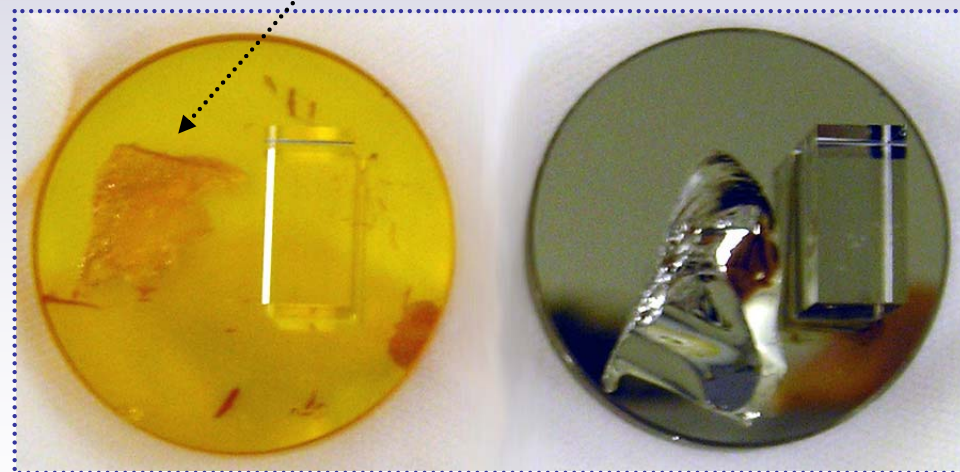


Temperature cycling:
 SiO_2 (silica) bonded to
 ZnSe (zinc selenide)
 and Ge (germanium)
 down to 77K



Images from video of temperature cycling
 of SiO_2 bonded to ZnSe down to 77K

Damage due mechanical strength testing.



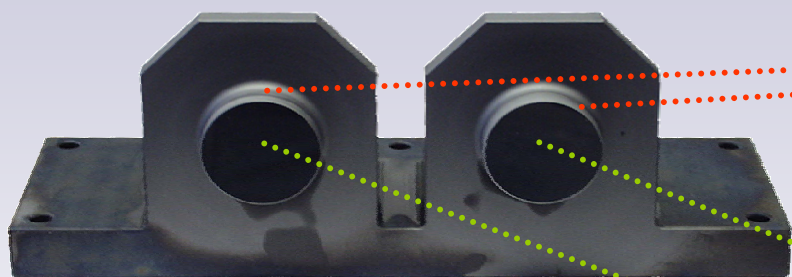
SiO_2 bonded
 to ZnSe

SiO_2 bonded
 to Ge



Investigations into hydroxy-catalysis bonding

- Astrium France – IGR subcontract – silicon carbide optical assemblies for GAIA – **patent application filed**
- **Current non-disclosure agreement in place** with TNO-TPD for further evaluation

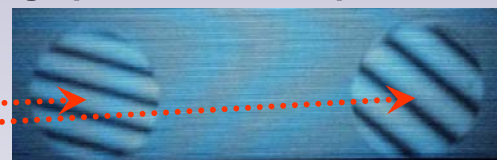


Picture of silicon carbide mirrors bonded to a silicon carbide base mount for GAIA

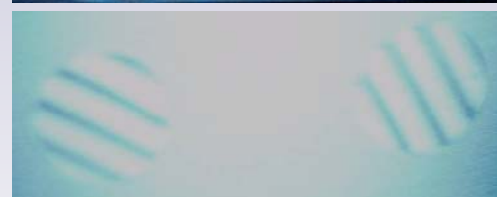
No relative tilt change between number of fringes within the resolution of the interferometer

Relative tilt in mirrors before and after bonding
< 1 arcsec (well within design specification of **3 arcsecs** in relative tilt).

fringe pattern of SiC optical mount

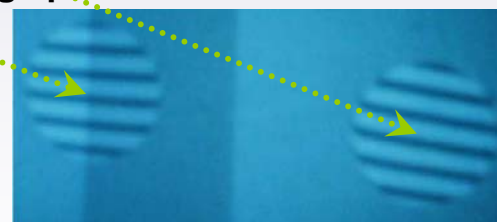


before oxidation



after oxidation

fringe pattern of mounted SiC mirrors



mirrors mounted dry



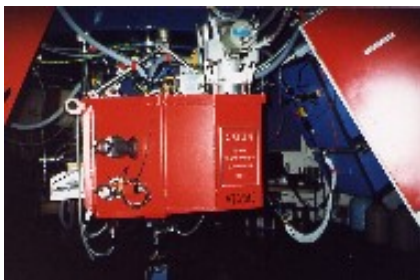
mirrors hydroxy-catalysis bonded

TEOPS Activity

- Submitted (Proposals)
 - European Gravitational Wave Observatory (EGO)
 - silica ribbons for monolithic suspensions - submitted
 - TEOPS Proposal part of larger PPARC proposal for ELT
 - Development of key passive & adaptive mirror technologies using SiC mirrors
 - ATLAS upgrade
 - Development of supermodule structures
- Discussions (Contract)
 - Advanced LIGO
 - Optical and Mechanical Systems
 - GEO 600 and Hannover AEI Prototype
 - Upgrades to existing detector
 - Isolation and Suspension design
- Other
 - Adam Smith Scholarships
 - Mech Eng & Chemistry
 - PIPSS & mini-PIPSS
 - British Consulate LA and Scottish Executive
 - Caltech / Glasgow (SUPA) expansion

Commonality

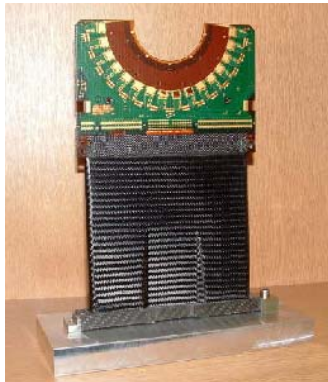
- Significant areas of commonality: some examples:
 - Cryogenics
 - The ATC has decades of experience in constructing reliable and robust instruments operating at cryogenic temperatures (as low as 4 K and even below 100 mK)
 - Cryogenic operation is now of interest for future generations of both gravitational wave detectors and colliders for particle physics



Commonality

■ New materials

- e.g. silicon-carbide being looked for use in astronomical instruments, particle physics detectors and in gravitational wave detectors
- CSiC trials underway in Glasgow
 - Thermal, mechanical, vacuum & bonding



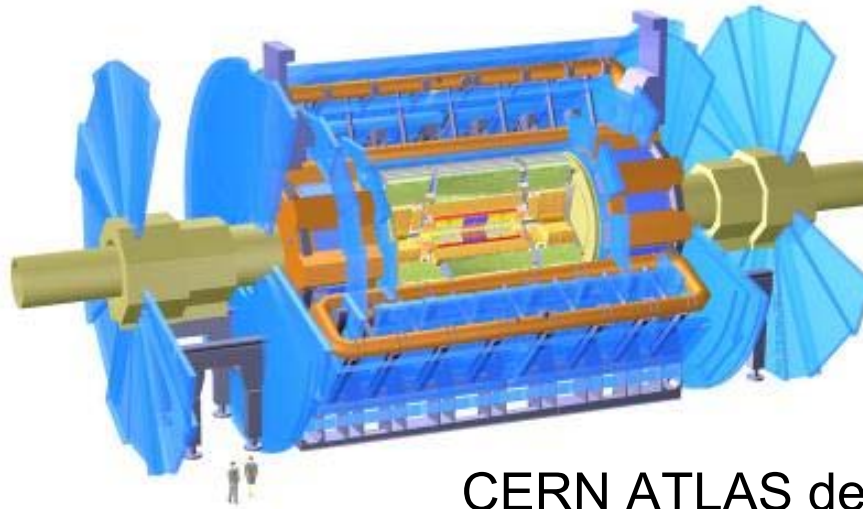
LHC module using carbon fibre mount
– Si-C considered as replacement



Si-C lightweighted telescope mirror
(courtesy M. Krodel)

Commonality

- Particle physics detector groups have experience in constructing detectors on an “industrial” scale
 - will be required in astronomy as telescopes increase in size (and number of telescopes in the case of arrays)



CERN ATLAS detector